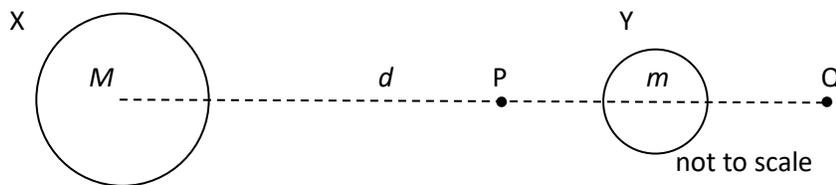


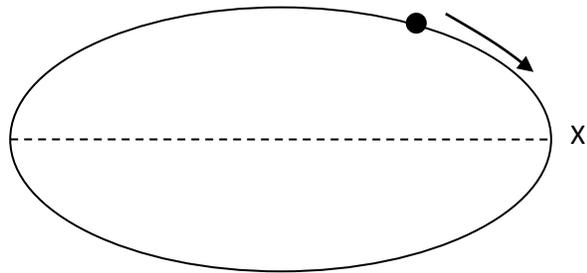
Problem of the week

Gravitational fields (SL)

- (a) The gravitational field strength on the surface of earth is g . What is the gravitational field strength
- at a distance of two earth radii from the **surface** of earth,
 - on the surface of a planet X that has double the mass and double the radius of earth,
 - on the surface of a planet Y that has the same density as earth and double the radius of earth.
- (b) In projectile motion we assume that the gravitational field above the earth's surface is uniform.
- Draw the gravitational field lines in this case.
 - Draw a realistic set of gravitational field lines for earth.
 - State what feature of the diagram in (ii) allows us to deduce that g decreases as we move away from the surface.
 - Suggest why two gravitational field lines cannot cross.
- (c) The diagram shows two planets, X and Y, of the same density and mass M and m . The planets are a distance d apart center to center.



- At point P, a distance $0.75d$ from the center of X, the gravitational field strength is zero. Determine the ratio $\frac{M}{m}$.
 - For the ratio in (i) determine the gravitational field strength at point Q, a distance of $0.25d$ to the right of the center of Y given that $\frac{GM}{d^2} = 2.0 \text{ N kg}^{-1}$.
 - Draw a graph (no numbers required) to show the variation of the resultant gravitational field with distance along the dotted line from the surface of X to the surface of Y. Take the positive direction to be that to the right.
- (d) The diagram shows the elliptical orbit of a planet around the Sun. As the planet approaches point X, the speed increases.



- (i) Draw the approximate position of the Sun,
- (ii) Explain your answer to (i).

(e) A planet is in a circular orbit of radius R around a star of mass M .

- (i) Show that the period of revolution of the planet is given by $T^2 = \frac{4\pi^2}{GM} R^3$.
- (ii) Io is a moon of Jupiter with an orbital radius of 4.2×10^8 m and an orbital period of 42 hours. Estimate the period of the Jupiter moon Callisto whose orbital radius is 1.9×10^9 m.

Answers

(a)

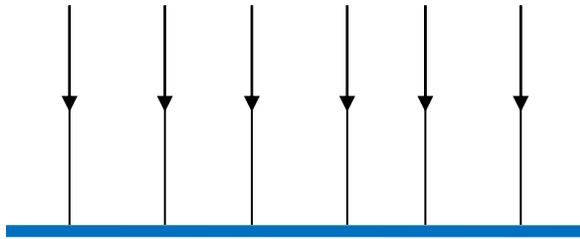
(i) $g = \frac{GM}{R^2}$ and $g' = \frac{GM}{(3R)^2} = \frac{1}{9} \frac{GM}{R^2} = \frac{g}{9}$.

(ii) $g' = \frac{G(2M)}{(2R)^2} = \frac{1}{2} \frac{GM}{R^2} = \frac{g}{2}$.

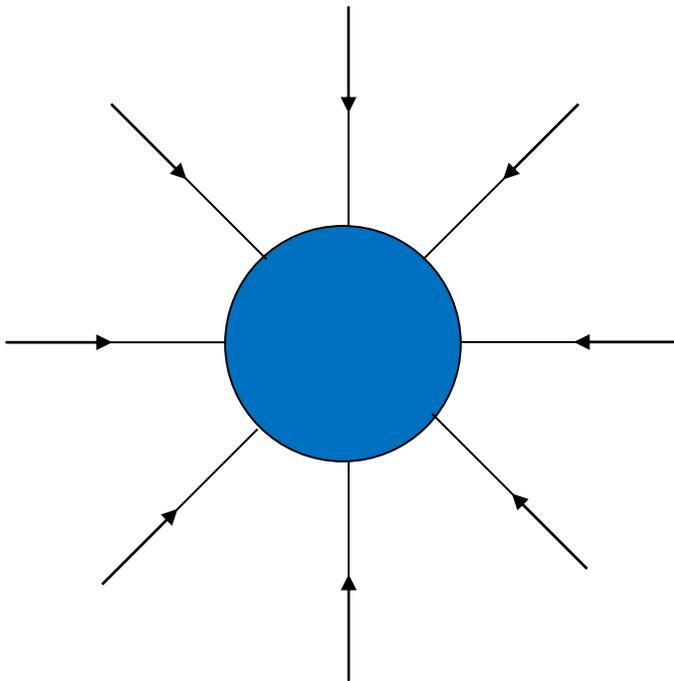
(iii) The volume is 8 times larger, so the mass is 8 times larger. Hence $g' = \frac{G(8M)}{(2R)^2} = 2 \frac{GM}{R^2} = 2g$.

(b)

(i)



(ii)



(iii) The field lines are getting further apart/their density is decreasing.

(iv) Tangents to field lines give the direction of the gravitational field which is unique at a point. Crossing lines would give two directions.

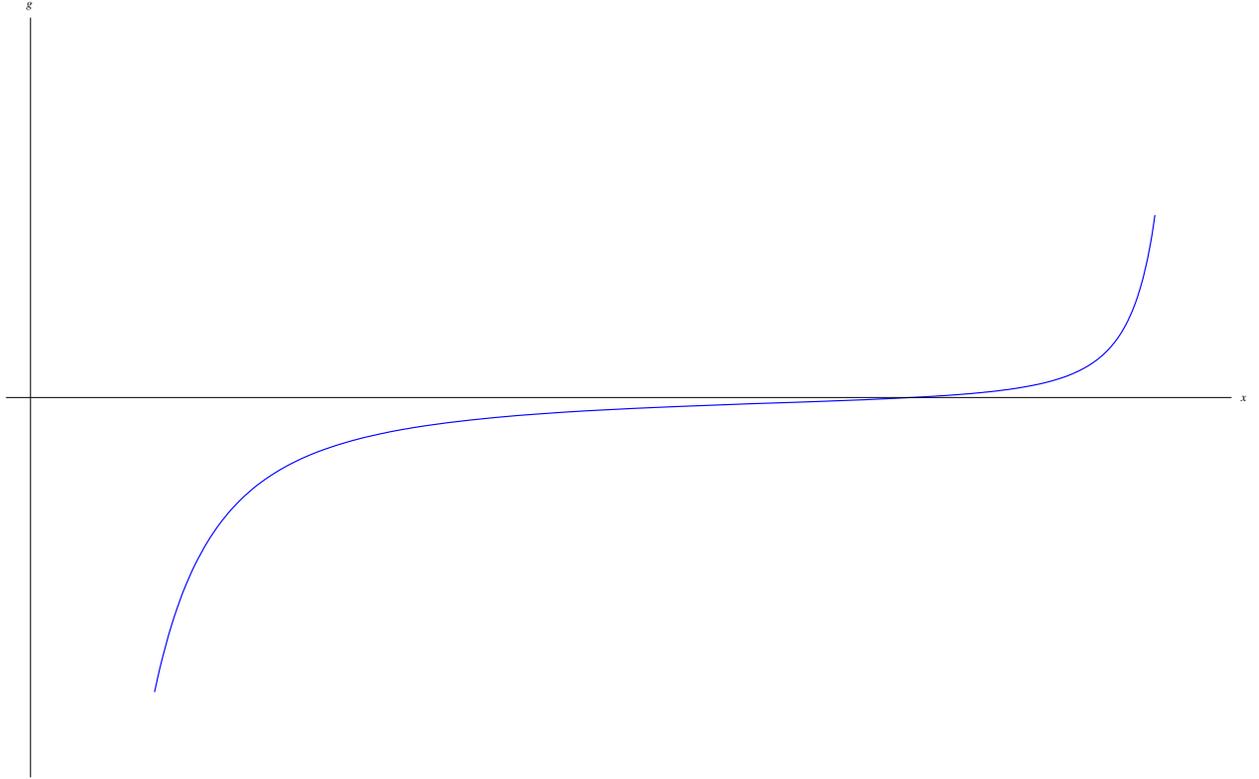
(c)

(i) The net field is zero so $\frac{GM}{(0.75d)^2} = \frac{Gm}{(0.25d)^2} \Rightarrow \frac{M}{m} = \left(\frac{0.75}{0.25}\right)^2 = 9.$

(ii) The net field at Q is

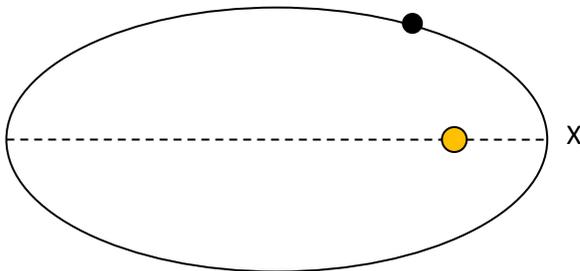
$$\frac{GM}{(1.25d)^2} + \frac{Gm}{(0.25d)^2} = \frac{GM}{(1.25d)^2} + \frac{GM}{9 \times (0.25d)^2} = \frac{544GM}{225d^2} = \frac{544}{225} \times 2.0 = 4.8 \text{ N kg}^{-1}.$$

(iii) Something like: (we care only about the general shape here)



(d)

(i)



(ii) There is a component of the gravitational force in the direction of velocity, so speed is increasing.

(e)

$$(i) \frac{GMm}{R^2} = \frac{mv^2}{R} \Rightarrow v^2 = \frac{GM}{R},$$

$$v = \frac{2\pi R}{T} \Rightarrow \left(\frac{2\pi R}{T}\right)^2 = \frac{GM}{R} \text{ hence, } T^2 = \frac{4\pi^2}{GM} R^3.$$

$$(ii) \frac{T_{\text{Callisto}}}{T_{\text{Io}}} = \left(\frac{R_{\text{Callisto}}}{R_{\text{Io}}}\right)^{\frac{3}{2}} = \left(\frac{1.9 \times 10^9}{4.2 \times 10^8}\right)^{\frac{3}{2}} = 96.2. \text{ Hence}$$

$$T_{\text{Callisto}} = 96.2 \times 42 = 4.04 \times 10^3 \text{ hr} = 16.8 \approx 17 \text{ days}$$